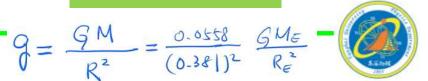
# N1-L08

## GPN1-L08





### Problem 1

Answer the following questions: (04小題)

$$= 0.381 \times 9.8 = 3.767$$

(a) Venus has a mass of about 0.0558 times the mass of Earth and a diameter of about 0.381 times the diameter of Earth. The acceleration of a body falling near the surface of Venus =  $m/s^2$ 

### 01: ANS:=3.8

(b) The mass of a planet is 1/100 that of Earth and its radius is 1/4 that of Earth. If a person weighs 600N on Earth, what would he weigh on this planet? W=\_\_\_\_N  $g = \frac{1}{(1)^2} \times 9.8 = 1.568$  W= mg =  $\frac{600}{9.8} \times 1.568 = 96$ 

(c) An object at the surface of Earth (at a distance R from the center of Earth) weighs 90N. Its weight at a distance 3R from the center of Earth = \_\_\_\_ N for r>R ⇒ Fx +2 90. == 10

(d) The approximate value of g at an altitude above Earth equal to one Earth diameter =  $m/s^2$ 

$$\gamma = R + D = R + 2R = 3R$$
  
 $g \propto \frac{1}{12} \Rightarrow g = \frac{9.8}{3^2} = 1.689$ 

A spherical shell has inner radius R1, outer radius R2, and mass M, distributed uniformly throughout the shell. Find the magnitude F of the gravitational force exerted on the shell by a point mass particle of m, located a distance d from the center for the following 3 cases. (03小題)

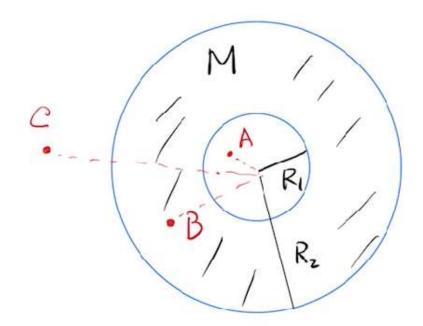
(a)
$$d < R1$$
,  $F =$ \_\_\_\_[M,R1,R2,G,d]

05: ANS:=0

(b)
$$R1 < d < R2, F = ____ [M,R1,R2,G,d]$$

06: ANS:=
$$\frac{G*M*m*(d**3-R1**3)}{(R2**3-R1**3)*(1/d**2)}$$

(c)
$$d>R2$$
,  $F=$ \_\_\_\_[M,R1,R2,G,d]



# 殼層定理(Shell Theorem)

1.球對稱物體對於球體外的重力 貢獻如同將球體質量集中於球心。 2.在對稱球體內部的物體不受其 外部球殼的重力影響。

A: 
$$d < R_1 \Rightarrow g = 0$$
,  $F = 0$ 

B:  $R < d < R_2 \Rightarrow M' = \alpha M$ ,  $M = w$  asso in side  $d$ 

$$F = \frac{GM'm}{r^2} \qquad \alpha = \frac{d^3 - R_1^3}{R_2^3 - R_1^3}$$

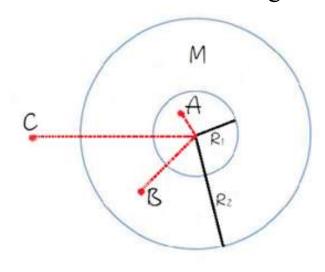
$$= \frac{G(\frac{d^3 - R_1^3}{R_2^3 - R_1^3})MM}{d^2}$$

$$= \frac{GMm}{d^2} \frac{d^3 - R_1^3}{R_2^3 - R_1^3}$$

C:  $d > R_2 \Rightarrow F = \frac{GMm}{d^2}$ 

# GPN1-L08/P02

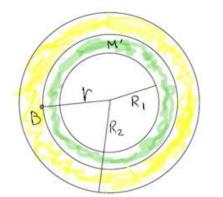
A spherical shell has inner radius R1, outer radius R2, and mass M, distributed uniformly throughout the shell. Find the magnitude FF of the gravitational force exerted on the shell by a point mass particle of m, located a distance d from the center for the following 3 cases.



教授不好意思 看完您的影片後 我還是有點不懂 PR06的解是如何求出的 能否請您更詳細説明呢?

B:R(-d < R<sub>2</sub> ⇒ M = d M (M= wass incide d)

B矣處的重力場是由下图中 总品色区域的質量所真截,外圍的 美色区域没有真南大 (Shell theorem). 所 以我們要先算出為色区的質量(M') 依据質量的体镜地,我們设H=XM M=輕片殼會的總質量(黃十絲) M'<M, X<1 M' >V'  $V' = \frac{4}{3}\pi r^3 - \frac{4}{3}\pi R_i^3$   $M \to V$  $V = \frac{4}{3}\pi R_2^3 - \frac{4}{3}\pi R_1^3$  $\frac{V'}{V} = \frac{\gamma^3 - R_1^3}{R_1^3 - R_1^3} = \frac{M'}{M} = \infty$ 



**殼層定理(Shell Theorem)**是古典重力學上的理論,其可簡化重力於對稱球體內部和外部的貢獻,並且在天文學上有特別的應用。 殼層定理最先由牛頓在所推演出來,其闡明了

- 1.球對稱物體對於球體外的重力貢獻如同將球體質量集中於球心。
- 2.在對稱球體內部的物體不受其外部球殼的重力影響。

由殼層定理的結果亦可得知,在一質量均勻分布的球體,重力由表面至中心線性遞減至零。因為球殼不會對內部物體有重力之貢獻,而剩餘之質量(不包括球殼)是與 $r^3$ 成正比,而重力是正比於 $m/r^2$ ,因此重力與 $r^3/r^2 = r$ 成正比。

Answer the following questions: (03小題)

(a)Each of the four corners of a square with edge a is occupied by a point mass m. There is a fifth mass, also m, at the center of the square. To remove the mass from the center to a point far away, find the work that must be done by an external agent.

the work=\_\_\_\_[G,m,a]

08: ANS:=4\*sqrt(2)\*G\*m\*\*2/a

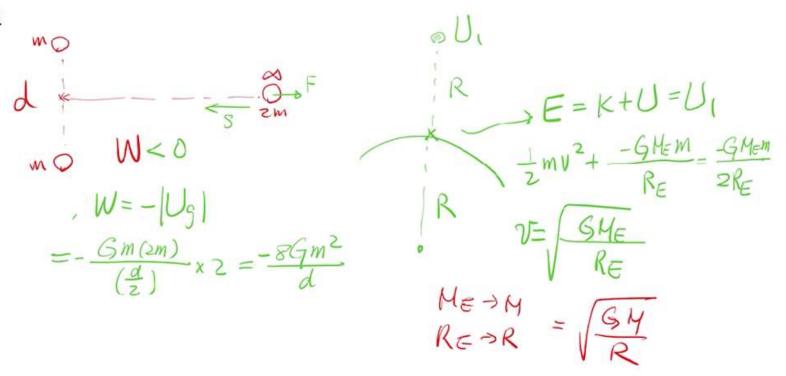
$$|\nabla u| = |\nabla u$$

(b)Two particles, each of mass m, are a distance d apart. To bring a third particle, with mass 2m, from far away to a resting point midway between the two particles the work done by an external agent=\_\_\_\_ [G,m,d]

### 09: ANS:=-8\*G\*m\*\*2/d\*\*2

(c) An object is dropped from an altitude of one Earth radius above Earth's surface. If M is the mass of Earth and R is its radius the speed of the object just before it hits Earth is given by, v =\_\_\_\_\_ [G,M,R]

### 10: ANS:=sqrt(G\*M/R)



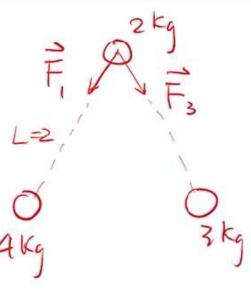
Three point particles with masses  $m_1 = 4 \text{ kg}$ ,  $m_2 = 2 \text{ kg}$ , and  $m_3 = 3 \text{ kg}$  are at the corners of an equilateral triangle of side L = 2 m, as in the figure. Find the net force  $\vec{F} = F_x \hat{i} + F_y \hat{j}$  on  $m_2$ . (02/1\Big)



11: ANS:=-1.6E-11

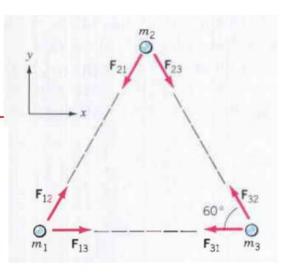
$$F_y$$
=\_\_\_\_N

12: ANS:=-20.3E-11



$$F_1 = \frac{G(4)(2)}{2^2} = \frac{8G}{4}$$

$$F_3 = \frac{GG}{2^2} = 1.5G$$



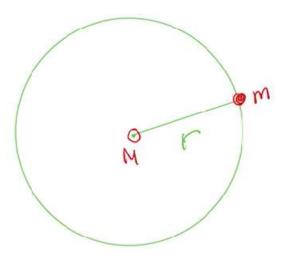
$$\frac{1}{F_{3}} = -\frac{1}{2}F_{1}^{2} - \frac{1}{2}F_{1}^{2} - \frac{1}{2}F_{1}^{$$

A satellite of mass m is in stable circular orbit of radius r. Find the following quantities: (a) the speed; (b) the period; (c) the linear momentum; (d) the kinetic energy; (e) the angular momentum? The mass of

Earth is *M*. (05小題)

16: ANS:=
$$(G*M*m)/(2*r)$$

(e) the angular momentum=\_\_\_\_ [G,M,m,r]



$$\frac{mr^2}{r} = \frac{GMm}{r^2}$$
,  $r = \sqrt{\frac{GM}{r}}$ 

$$T = \frac{2\pi r}{v} = \frac{2\pi r}{\sqrt{\frac{GM}{r}}} = 2\pi \sqrt{\frac{r^3}{GM}}$$

$$P = mV = m \sqrt{\frac{GM}{r}}$$

$$K = \frac{1}{2} mV^2 = \frac{1}{2} m \left(\frac{GM}{r}\right) = \frac{GMm}{2r}$$

**殼層定理(Shell Theorem)**是古典重力學上的理論,其可簡化重力於對稱球體內部和外部的貢獻,並且在天文學上有特別的應用。 殼層定理最先由牛頓在所推演出來,其闡明了

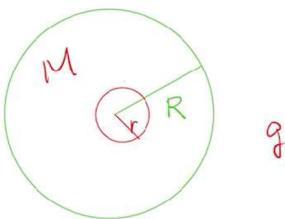
- 1.球對稱物體對於球體外的重力貢獻如同將球體質量集中於球心。
- 2.在對稱球體內部的物體不受其外部球殼的重力影響。

由殼層定理的結果亦可得知,在一質量均勻分布的球體,重力由表面至中心線性遞減至零。因為球殼不會對內部物體有重力之貢獻,而剩餘之質量(不包括球殼)是與 $r^3$ 成正比,而重力是正比於 $m/r^2$ ,因此重力與 $r^3/r^2 = r$ 成正比。

How does the field strength g(r) vary inside a uniform solid sphere of density  $\rho$ =rho and radius R? (01/力) 題)

$$g(r)$$
=\_\_\_\_[G,rho,R,r]

**殼層定理(Shell Theorem)** 在對稱球體內部的物體不 受其外部球殼的重力影響。



$$M(r) = \frac{r^{3}}{R^{3}}M, M = \frac{4}{3}\pi R^{3}\rho$$

$$g(r) = \frac{GM(r)}{r^{2}} = \frac{GM}{R^{3}}r$$

$$= \frac{4}{3}\pi\rho Gr$$

#### Problem 6

Calculate the gravitational field strength at the surface of the following bodies (assumed to be uniform spheres): (a) Jupiter  $(1.9 \times 10^{27} \text{ kg}, \text{ radius } 7.14 \times 10^7 \text{ m})$ ; (b) a neutron star of mass  $10^{30}$  kg and radius 20 km.  $(02 \text{ /} \text{$ 

(a)
$$g_{\text{Jupiter}} =$$
\_\_\_\_\_ m/s<sup>2</sup>

(b) for neutron star,  $g = ___ m/s^2$ 

$$g = \frac{GM}{R^2}$$

$$g_{N} = \frac{6.67 \times 10^{11} \times 1.9 \times 10^{27}}{(7.44 \times 10^{27})^{2}} = 24.9$$

$$g_{N} = \frac{6.67 \times 10^{11} \times 10^{20}}{(20 \times 10^{3})^{2}} = 1.668 \times 10^{11} \text{ m/s}_{52}$$

In its elliptical orbit, the speed of the earth at perihelion(遠日點) is  $v_P = 3.03 \times 10^4$  m/s. If the distances to the sun at perihelion and aphelion(近日點) are  $r_P = 1.47 \times 10^{11}$  m and  $r_A = 1.52 \times 10^{11}$  m, find  $v_A$ . (01小題)

$$v_A$$
=\_\_\_\_m/s

21: ANS:=2.93E4

$$Lp = LA$$

$$r_{p} \nabla_{p} = r_{A} \nabla_{A}$$

$$(1.47 \times 10^{11})(3,03 \times 10^{4}) = (1.52 \times 10^{11}) V_{A}$$

$$V_{A} = 2.93 \times 10^{4}$$